

# Shaul's List

What would be the impact of a positive B-mode detection?

What would be the impact of no B-mode detection at a level of  $r=0.01$ ?

What are the arguments that  $r$  is larger than 0.01?

What would be the value of a mission that measured lensing B modes very well, but did not detect primordial B modes?

What priority would you give a \$1B investment in a CMB polarization mission compared to other possible astrophysics space missions?

If you think it is worthwhile, how would you articulate this to colleagues?

What are the arguments that  $r > 0.01$  ?

## Inflation

generically produces a nearly scale-invariant spectrum of gravitational waves with amplitude roughly comparable to the density fluctuation amplitude

## Alternatives

the only known alternatives for producing adiabatic, nearly scale-invariant perturbations – using the ekpyrotic mechanism – do not

What are the arguments that  $r > 0.01$  ?

## Inflation

fluctuations are due to a gravitational effect of the background that is indiscriminant: all light degrees of freedom are subject to the same fluctuations.

## Alternatives

primordial fluctuations are due to a non-gravitational effect that only affects fields with steep, negative potentials (therefore, not g-waves)

What are the arguments that  $r > 0.01$  ?

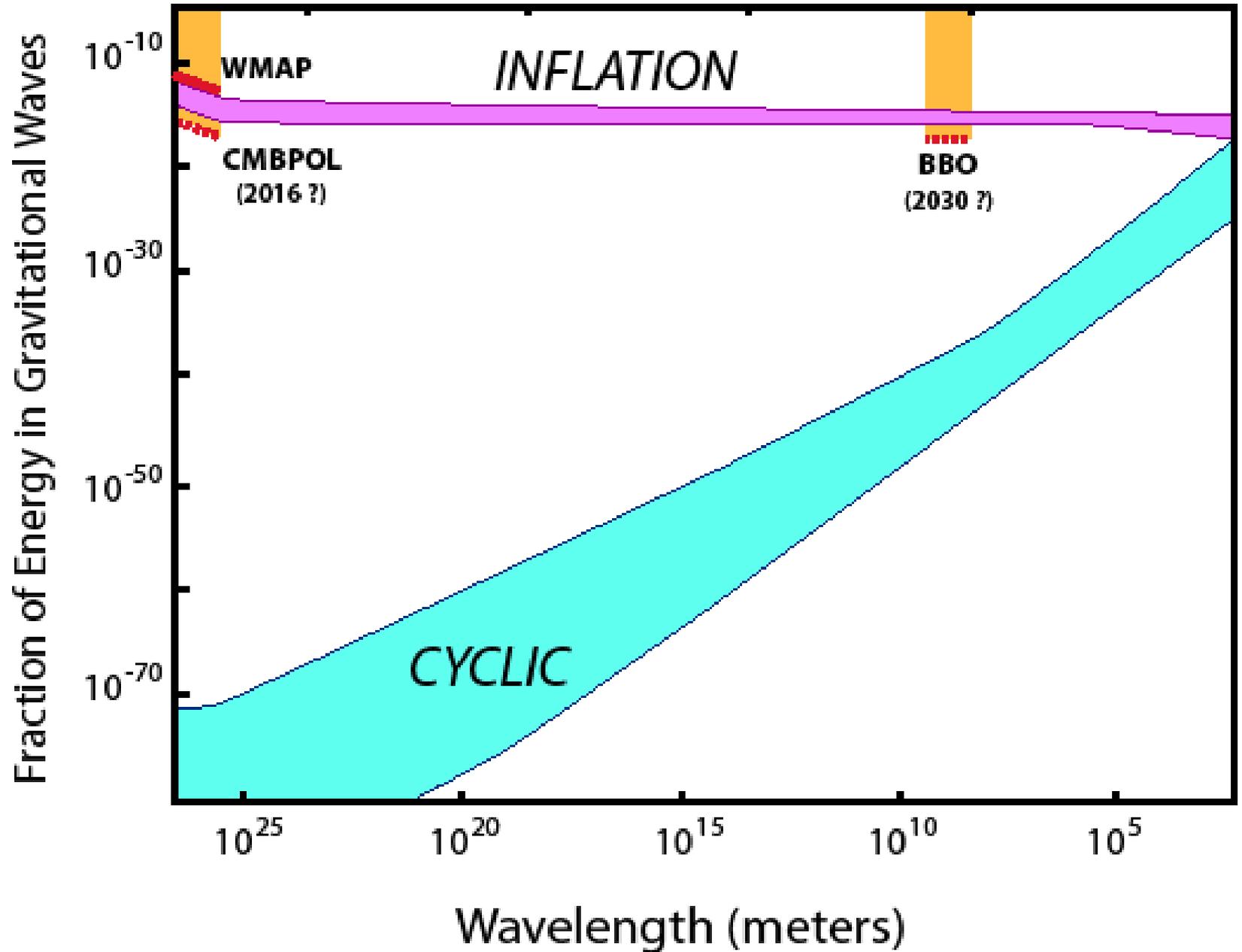
## Inflation

$$r \sim 0.01 - 0.4^*$$

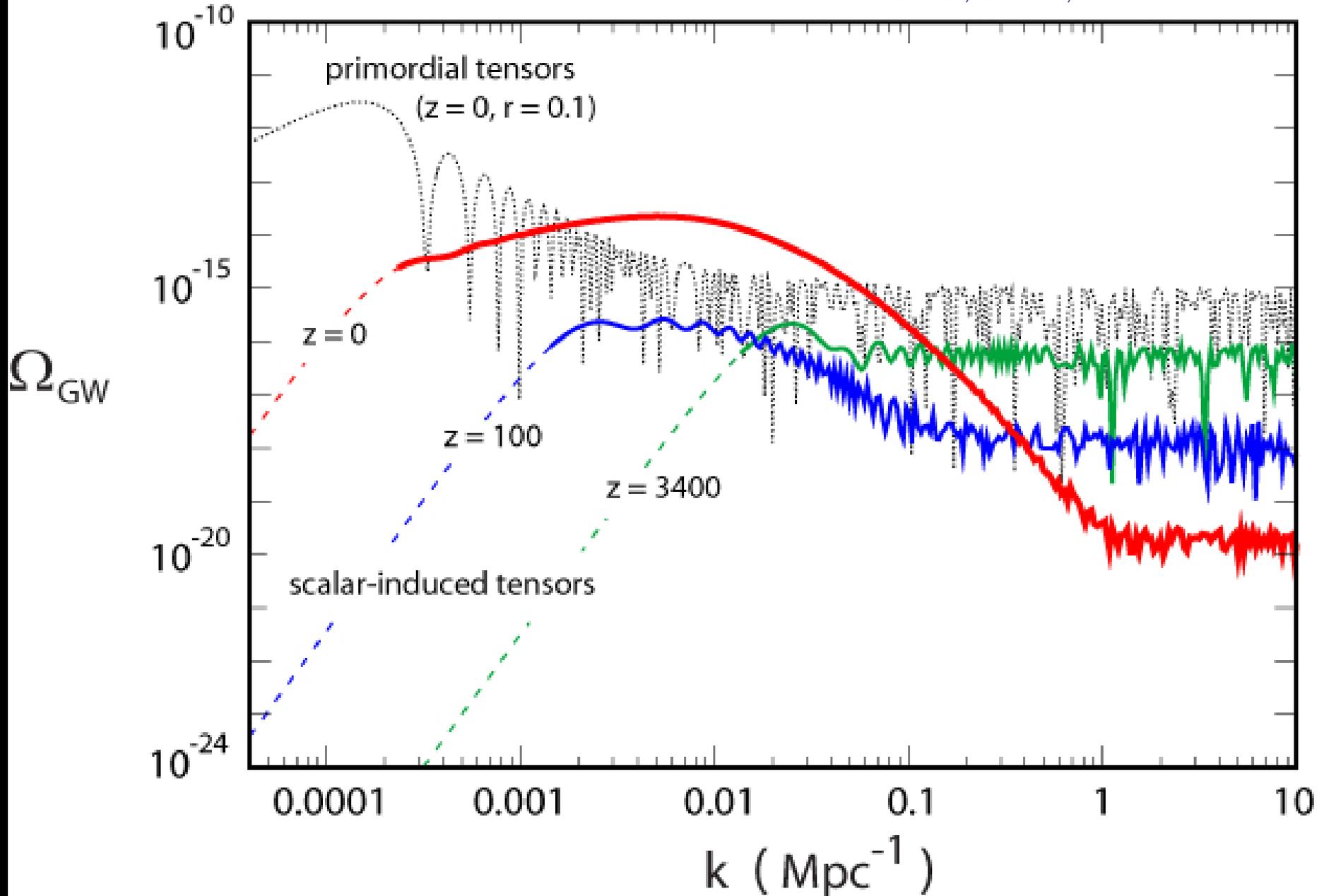
## Alternatives

primordial  $r$  exponentially suppressed

# primordial g-waves



from Baumann, PJS, Takahashi, Ichiki  
see also Mollerach, Harari, Mattarese  
Ananda, Clarkson, Wands



What are the arguments that  $r > 0.01$  ?

## Inflation

$$r \sim 0.01 - 0.4^*$$

\*qualifications

$$r \sim (1+w)^{1/4}$$

## Alternatives

primordial  $r$  exponentially suppressed

# What would be the impact of a positive B-mode detection?

- opens the window to measuring the g-wave spectrum (e.g., tilt)
- verifies the natural prediction of inflation
- determines energy scale for physics beyond the standard model that created large scale structure of the universe
- supports notion that fluctuations and seeds for galaxy formation were created by events occurring after the big bang
- lends support to the big bang being the effective beginning of spacetime
- definitively eliminates the competing ekpyrotic/cyclic alternatives
- definitively eliminates (almost) all current string inflation models

# What would be the impact of no B-mode detection at a level $r > 0.01$ ?

- first (?) serious blow to simplest inflation models
- restricts future consideration of inflation to models with extra degrees of freedom and finely-tuned parameters
- verifies the natural prediction of ekpyrotic/cyclic models
- proves that the fluctuations were created during a period of small  $H$ :
  - slow contraction (ekpyrotic)
  - slow expansion (finely-tuned inflation)
- sets upper bound for energy scale for new physics that created the large scale structure of the universe
- lends support to idea of fluctuations created before the bang, hence encourages research on big bang physics and novel solutions to other cosmological problems

# What would be the impact of no B-mode detection at a level $r > 0.01$ , but measurement of lensing B-modes?

- add confidence in the absence of primordial B-modes
- improve measurement of lensing, large scale structure, cosmological parameters
- test for cosmic strings
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